

Power Source for the Most Energetic Explosions in the Local Universe

Super-Luminous Supernova Gaia16apd

Posner Program



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Evolution of a Star



| | | | | | | | | | | | | | | | | | | | | |
|---------|----------|----------|----------|----------|----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|
| Group→ | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 | 11 | 12 | 13 | 14 | 15 | 16 | 17 | 18 | | |
| ↓Period | | | | | | | | | | | | | | | | | | | | |
| 1 | 1 H | | | | | | | | | | | | | | | | | | 2 He | |
| 2 | 3 Li | 4 Be | | | | | | | | | | | 5 B | 6 C | 7 N | 8 O | 9 F | 10 Ne | | |
| 3 | 11 Na | 12 Mg | | | | | | | | | | | 13 Al | 14 Si | 15 P | 16 S | 17 Cl | 18 Ar | | |
| 4 | 19 K | 20 Ca | 21 Sc | 22 Ti | 23 V | 24 Cr | 25 Mn | 26 Fe | 27 Co | 28 Ni | 29 Cu | 30 Zn | 31 Ga | 32 Ge | 33 As | 34 Se | 35 Br | 36 Kr | | |
| 5 | 37 Rb | 38 Sr | 39 Y | 40 Zr | 41 Nb | 42 Mo | 43 Tc | 44 Ru | 45 Rh | 46 Pd | 47 Ag | 48 Cd | 49 In | 50 Sn | 51 Sb | 52 Te | 53 I | 54 Xe | | |
| 6 | 55 Cs | 56 Ba | 57 La | * | 72 Hf | 73 Ta | 74 W | 75 Re | 76 Os | 77 Ir | 78 Pt | 79 Au | 80 Hg | 81 Tl | 82 Pb | 83 Bi | 84 Po | 85 At | 86 Rn | |
| 7 | 87 Fr | 88 Ra | 89 Ac | ** | ** | 104 Rf | 105 Db | 106 Sg | 107 Bh | 108 Hs | 109 Mt | 110 Ds | 111 Rg | 112 Cn | 113 Nh | 114 Fl | 115 Mc | 116 Lv | 117 Ts | 118 Og |
| | | | | | * | 58 Ce | 59 Pr | 60 Nd | 61 Pm | 62 Sm | 63 Eu | 64 Gd | 65 Tb | 66 Dy | 67 Ho | 68 Er | 69 Tm | 70 Yb | 71 Lu | |
| | | | | | ** | 90 Th | 91 Pa | 92 U | 93 Np | 94 Pu | 95 Am | 96 Cm | 97 Bk | 98 Cf | 99 Es | 100 Fm | 101 Md | 102 No | 103 Lr | |

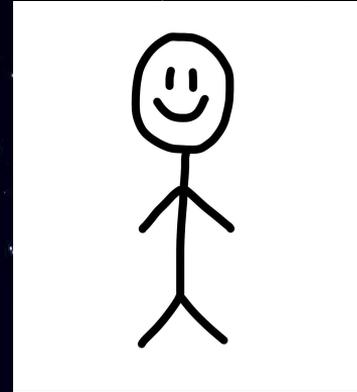


Explosion!!!

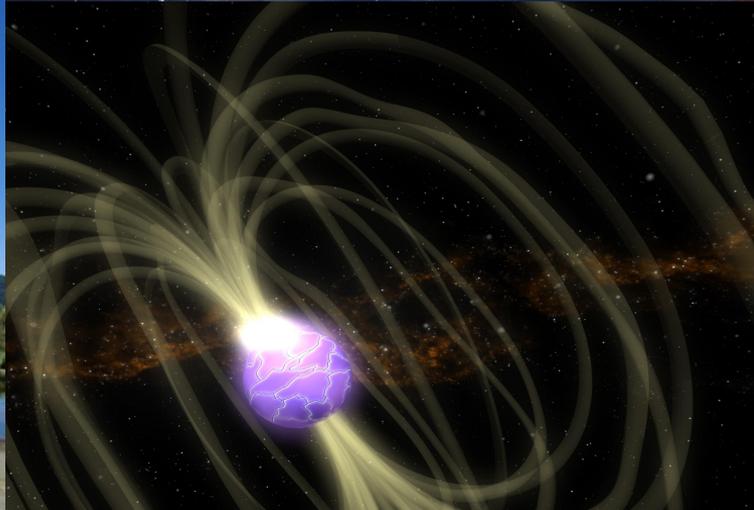


The Big Question:
What Powers this explosion?

But First - Why Does this Matter?



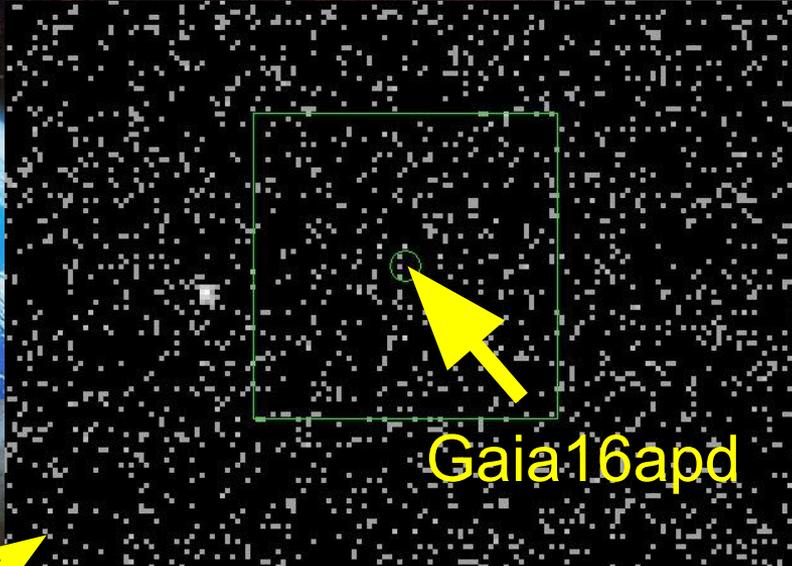
The Magnetar: The Best of the Three Models Explaining the Power Source



X-Ray Telescopes and Data



XMM Newton



Gaia16apd

X-Ray Photons
From Chandra

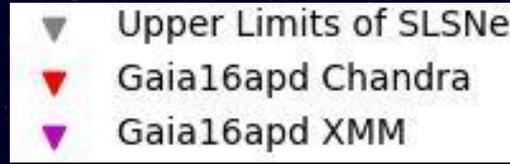


Chandra

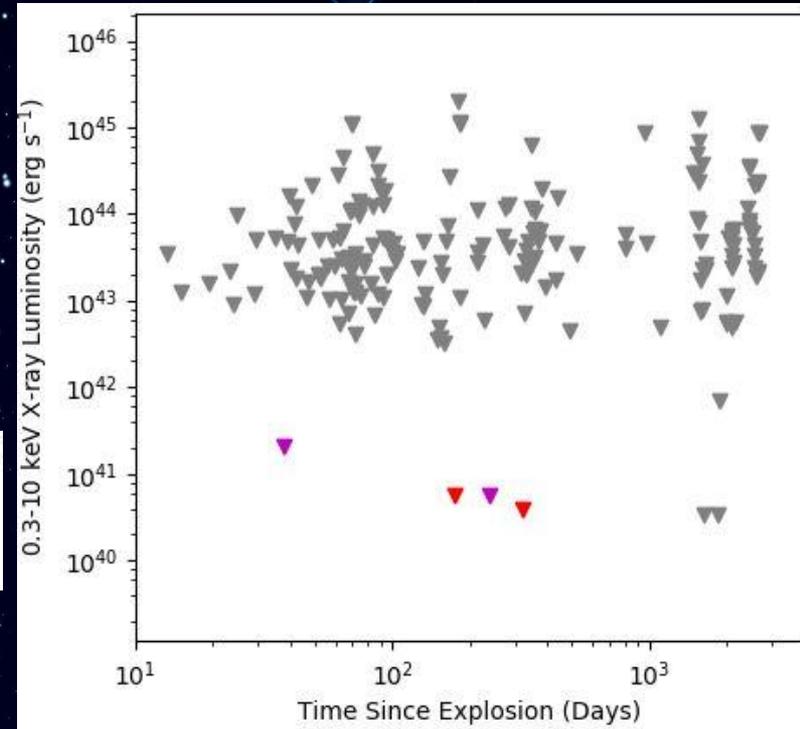
We used 4 of these
observations

Parameters for Jet Simulations

- Kinetic Energy
- Circumstellar Material Density
- Jet Opening Angle
- Observer Angle
- Epsilon E
- Epsilon B

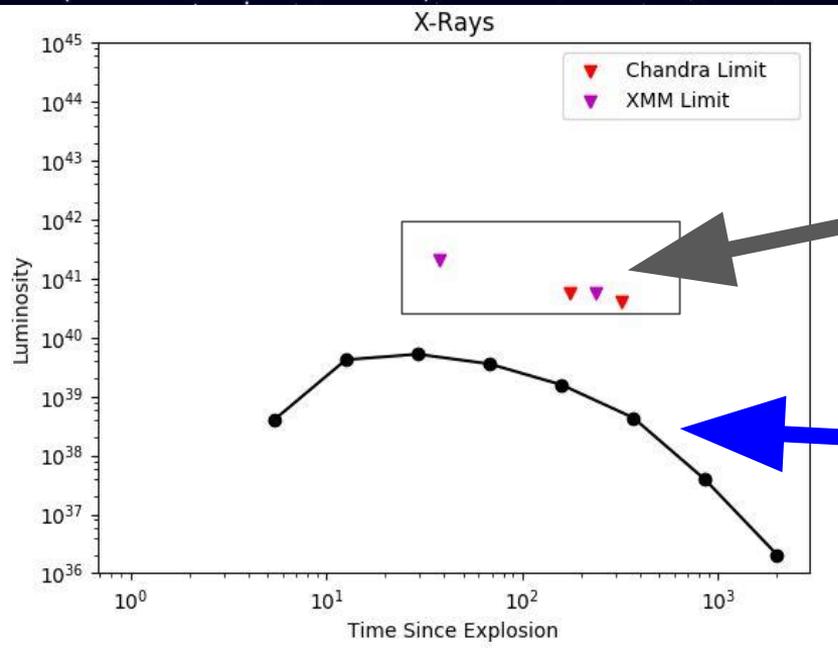


X-Ray Limits

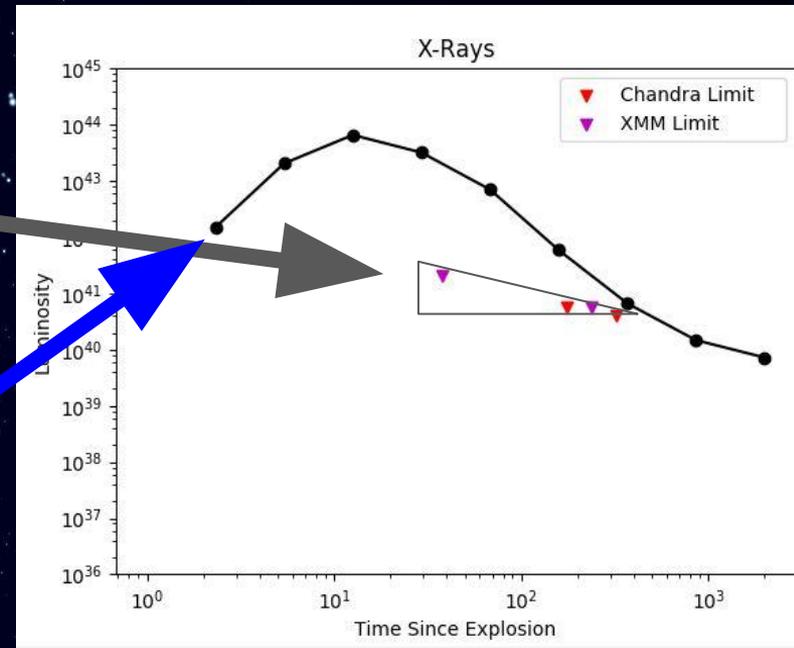


700+ Jet Simulations

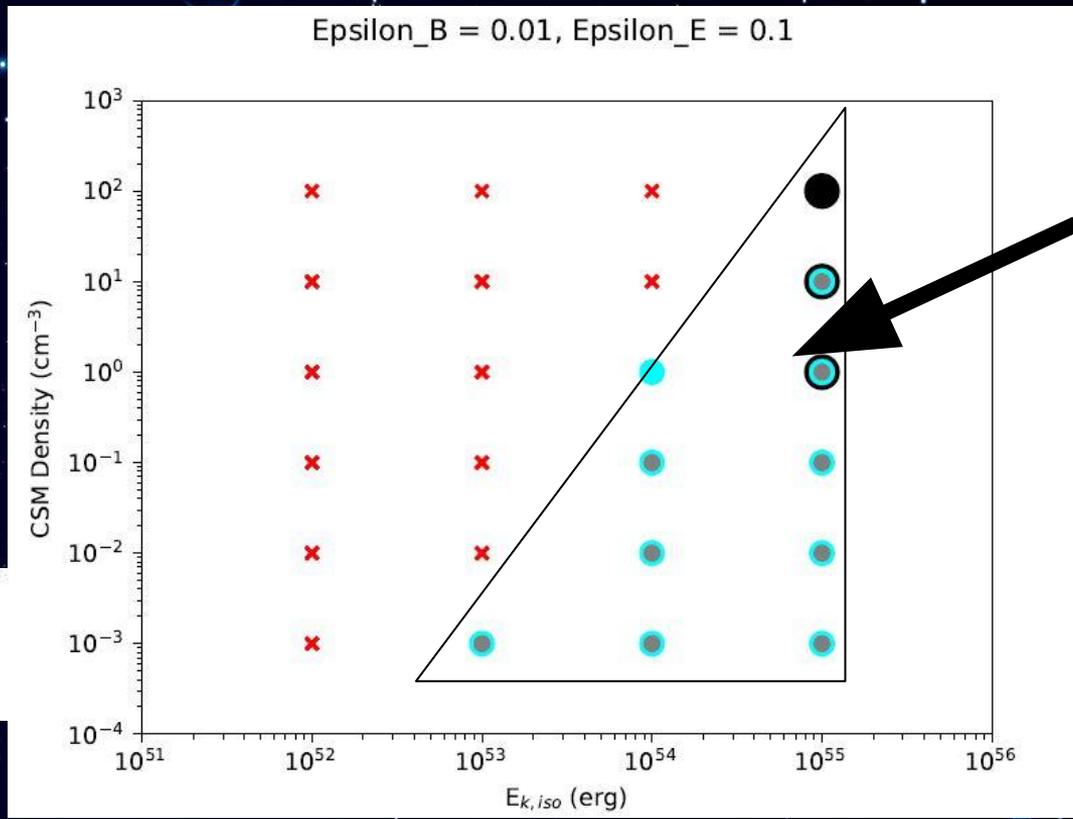
This Jet is plausible



This Jet is Ruled Out



What We Learned From This



Ruled out

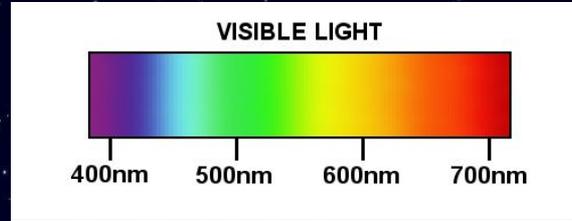
- Red X: Obs Not Ruled out for Jet $5^\circ, 10^\circ, 15^\circ$
- Black Circle: Jet = 5° , Obs = 30°
- Grey Circle: Jet = 5° , Obs = 7.5°
- Cyan Circle: Jet = 5° , Obs = 10°

Conclusion

- Observations have ruled out environments with high kinetic energy
- This means that the most powerful jets are not behind the most energetic explosions



The Electromagnetic Spectrum



Gamma Ray

X-Ray

Ultraviolet

Visible

Infrared

Microwave

Radio Wave

Atom Radius

0.01 - 10 nm

Hair Width

